#### UNIVERSITY OF CAMBRIDGE INTERNATIONAL EXAMINATIONS

GCE Advanced Subsidiary Level and GCE Advanced Level

# MARK SCHEME for the October/November 2010 question paper for the guidance of teachers

## 9702 PHYSICS

9702/42

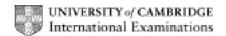
Paper 4 (A2 Structured Questions), maximum raw mark 100

This mark scheme is published as an aid to teachers and candidates, to indicate the requirements of the examination. It shows the basis on which Examiners were instructed to award marks. It does not indicate the details of the discussions that took place at an Examiners' meeting before marking began, which would have considered the acceptability of alternative answers.

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## **Section A**

1	(a) force per unit mass	(ratio idea essential)	B1 [1]
•	(a) Toroo per arm made	(ratio rada decertiar)	י פ

(b) graph: correct curvature M1 from 
$$(R,1.0\,g_{\rm S})$$
 & at least one other correct point A1 [2]

(c) (i) fields of Earth and Moon are in opposite directions

either resultant field found by subtraction of the field strength

or any other sensible comment A1

so there is a point where it is zero

(allow 
$$F_E = -F_M$$
 for 2 marks)

(ii) 
$$GM_E/x^2 = GM_M/(D-x)^2$$
 C1  
 $(6.0 \times 10^{24})/(7.4 \times 10^{22}) = x^2/(60R_E-x)^2$  C1  
 $x = 54R_E$  A1 [3]

(iii) graph: 
$$g = 0$$
 at least  $\frac{2}{3}$  distance to Moon B1  $g_{\rm E}$  and  $g_{\rm M}$  in opposite directions M1 correct curvature (by eye) and  $g_{\rm E} > g_{\rm M}$  at surface A1 [3]

- 2 (a) (i) no forces (of attraction or repulsion) between atoms / molecules / particles B1 [1]
  - (ii) sum of kinetic and potential energy of atoms / molecules M1 due to random motion A1 [2]
  - (iii) (random) kinetic energy increases with temperature no potential energy (so increase in temperature increases internal energy)

    A1 [2]
  - (b) (i) zero A1 [1]

(ii) work done = 
$$p\Delta V$$
 C1  
=  $4.0 \times 10^5 \times 6 \times 10^{-4}$   
= 240 J (ignore any sign) A1 [2]

(iii)

change	work done / J	heating / J	increase in internal energy / J
$\begin{array}{c} P \rightarrow Q \\ Q \rightarrow R \\ R \rightarrow P \end{array}$	<b>+240</b> 0 <b>-840</b>	-600 +720 +480	-360 +720 -360

(correct signs essential)
(each horizontal line correct, 1 mark – max 3)

B3 [3]

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3	(a)	(i)	reson	ance		B1	[1]
		(ii)	amplit	tude 16 mm <u>and</u> frequency 4.6 Hz		A1	[1]
	(b)	(i)	a = 4	$-)\omega^2 x$ and $\omega = 2\pi f$ $4\pi^2 \times 4.6^2 \times 16 \times 10^{-3}$ $43.4 \mathrm{m  s^{-2}}$		C1 C1 A1	[3]
		(ii)	F = n = 1	ma ∣50 × 10 <sup>−3</sup> × 13.4		C1	
			= 2	2.0 N		A1	[2]
	(c)			s 'below' given line and never zero 4.6 Hz (or slightly less) and flatter		M1 A1	[2]
4	(a)	cha	rge / p	otential (difference) (ratio must be clear)		B1	[1]
	(b)	(i)	V = Q	$0/4\pi\varepsilon_0 r$		B1	[1]
		(ii)	C = Q so $C < Q$	$0/V = 4\pi \varepsilon_0 r$ and $4\pi \varepsilon_0$ is constant $\infty r$		M1 A0	[1]
	(c)	(i)	r = C ; r = (6. = 6.1	/ $4\pi\varepsilon_0 r$ 8 × $10^{-12}$ ) / $(4\pi \times 8.85 \times 10^{-12})$ × $10^{-2}$ m		C1 C1 A1	[3]
		(ii)		$EV = 6.8 \times 10^{-12} \times 220$ $1.5 \times 10^{-9}$ C		A1	[1]
	(d)	(i)	V = Q = 83 \	$V/C = (1.5 \times 10^{-9}) / (18 \times 10^{-12})$		A1	[1]
		(ii)	either	$\Delta E = \frac{1}{2} \times 6.8 \times 10^{-12} \times 220^2 - \frac{1}{2} \times 18 \times 10^{-12} \times 83$	2	C1 C1	
			or	= $1.65 \times 10^{-7} - 6.2 \times 10^{-8}$ = $1.03 \times 10^{-7}$ J energy = $\frac{1}{2}$ QV $\Delta E = \frac{1}{2} \times 1.5 \times 10^{-9} \times 220 - \frac{1}{2} \times 1.5 \times 10^{-9} \times 83$ = $1.03 \times 10^{-7}$ J		A1 (C1) (C1) (A1)	[3]

		9	GCE AS/A LEVEL – October/November 2010 9702	42	
5	(a)	field i	nto (the plane of) the paper	B1	[1]
	(b)	mv² / B =	due to magnetic field <u>provides</u> the centripetal force $r = Bqv$ = $(20 \times 1.66 \times 10^{-27} \times 1.40 \times 10^5) / (1.6 \times 10^{-19} \times 6.4 \times 10^{-2})$ = 0.454 T	B1 C1 B1 A0	[3]
	(c)	(i) <u>s</u>	semicircle with diameter greater than 12.8 cm	B1	[1]
		(ii) r	new flux density = $\frac{22}{20}$ × 0.454	C1	
			B = 0.499 T	A1	[2]
6	(a)	(i) e	e.g. prevent flux losses / improve flux linkage	B1	[1]
			lux in core is changing e.m.f. / current (induced) <u>in core</u> nduced current in core causes heating	B1 B1 B1	[3]
	(b)		hat value of the direct current producing same (mean) power / heating n a resistor	M1 A1	[2]
			power in primary = power in secondary $V_P I_P = V_S I_S$	M1 A1	[2]
7	(a)	(i) e	e.g. electron / particle diffraction	B1	[1]
		(ii) ∈	e.g. photoelectric effect	B1	[1]
	(b)	(i) 6		A1	[1]
		` ' /	change in energy = $4.57 \times 10^{-19}$ J R = hc / E $= (6.63 \times 10^{-34} \times 3.0 \times 10^{8}) / (4.57 \times 10^{-19})$	C1	
			$= 4.4 \times 10^{-7} \text{ m}$	A1	[2]
8	(a)		ng of a heavy nucleus (not atom/nuclide) wo (lighter) nuclei of approximately same mass	M1 A1	[2]
	(b)	$_{0}^{1}$ n $_{2}^{4}$ He $(allow\ _{2}^{4}\alpha\ )$ $_{3}^{7}$ Li		M2 A1	[3]
	(c)	emitted particles have kinetic energy range of particles in the control rods is short / particles stopped in rods /			
		lose l kineti	B1 B1	[3]	

Mark Scheme: Teachers' version

Syllabus

Paper

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## Section B

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	•	the n	nicrophone iers scores no mark)		M1 A1	[2]
2 (a)	satellite receives signal amplifies at a different (different frequere.g. of frequere.g.	es gre d and carrier encies ncies u	nitted from Earth to satellite eatly attenuated signal transmitted back to Earth ) frequency prevent swamping of uplink signal sed (6/4 GHz, 14/11 GHz, 30/20 GHz) any two other for additional physics)	(1) (1) (1) (1)	B1 B1	[4]
(b)	advantage:	e.g. e.g.	because orbits are much lower		M1 A1 (M1) (A1)	
	disadvantage:	e.g.	<ul><li>either must be tracked</li><li>or limited use in any one orbit</li><li>more satellites required for continuous of</li></ul>	pperation	M1 A1	[4]